

# Sonic black holes and Hawking radiation in Bose-Einstein condensates

Nicolas Pavloff

Laboratoire de Physique Théorique et Modèles Statistiques  
Université Paris-Sud, CNRS, Orsay



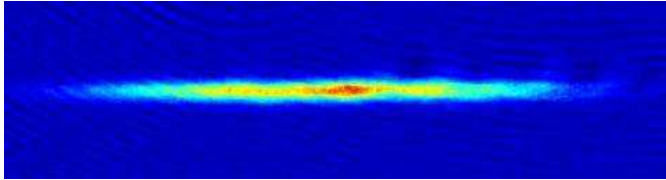
work in collaboration with:



I. Carusotto and A. Recati

P.É. Larré

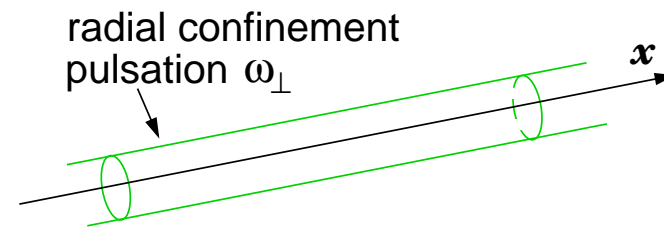
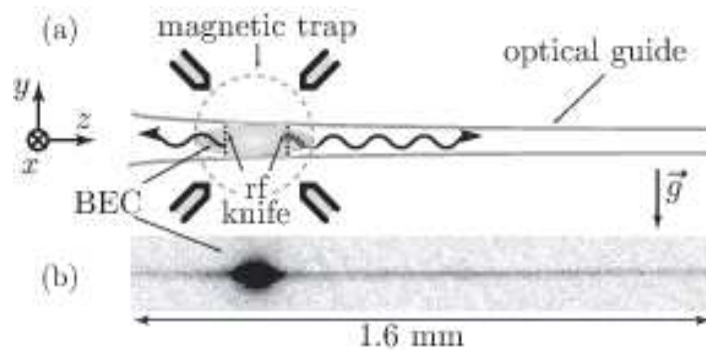
## quasi-1D condensates :



*quasi-1D condensate*

longitudinal size  $\sim 10^2 \mu\text{m}$

transverse size  $\sim 1 \mu\text{m}$



harmonic radial confinement :

$$V_{\perp}(\vec{r}_{\perp}) = \frac{1}{2} m \omega_{\perp}^2 r_{\perp}^2 .$$

W. Guérin *et al.*, Phys. Rev. Lett. **97**, 200402 (2006)

## Mesoscopic physics & BECs :

interaction in phase coherent systems,  
non-linear transport.

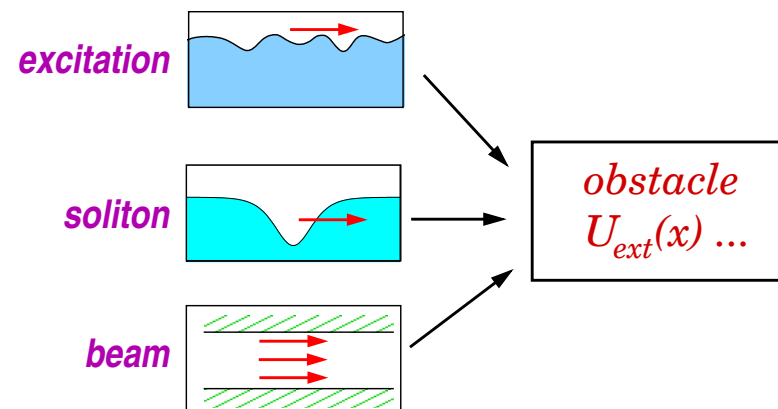
Large range of interaction regimes :

↪ From “atom lasers” practicaly without interaction → strongly correlated 1D systems

↪ well defined theoretical framework (Bose-Hubbard/Gross-Pitaevskii)

### Situations of 1D transport :

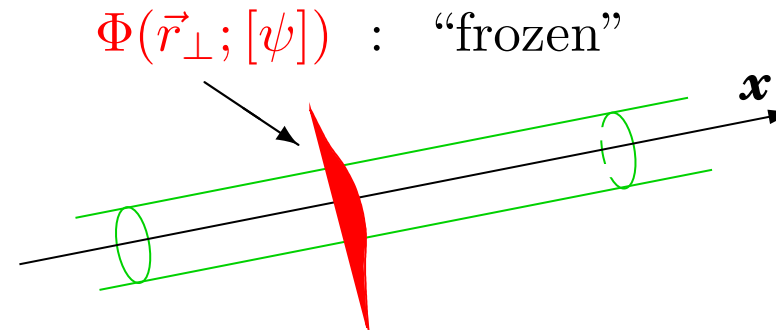
- Propagation of excitations, of (dark) solitons, of a beam ...
- In presence of localized or extended obstacles
- Effects of disorder
- Black-hole configuration
- Dispersive shock waves



## 1D mean field regime

Born-Oppenheimer :

$$\Psi(\vec{r}, t) = \psi(x, t) \times \Phi(\vec{r}_\perp; [\psi])$$



1D mean field regime with order parameter  $\psi(x, t)$  verifying

$$-\frac{\hbar^2}{2m} \partial_x^2 \psi + (U_{\text{ext}}(x) + g |\psi|^2) \psi = i\hbar \partial_t \psi \quad \text{or} \quad \mu \psi \quad (1)$$

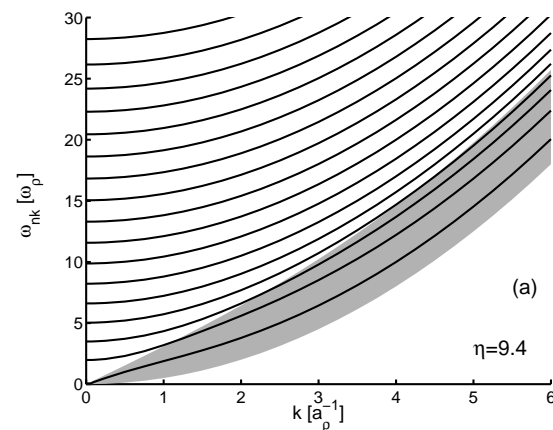
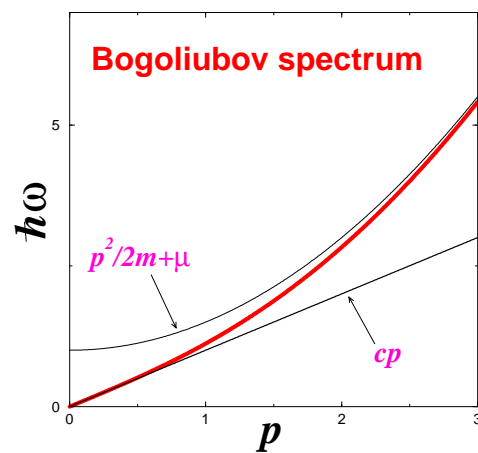
where  $|\psi|^2 = n_1(x, t)$  is the longitudinal density of the condensate,  
and  $g = 2\hbar\omega_\perp a$ , where  $a$  : 3D s-wave scattering length ( $a > 0$ )

domain of validity :

$$\frac{\hbar\omega_{\perp}}{\hbar^2/ma^2} \ll n_1 a \sim \frac{\mu}{\hbar\omega_{\perp}} \ll 1$$

• The first inequality allows to avoid the **Tonks-Girardeau regime** and implies  $E_{\text{int}} \ll E_{\text{kin}}$ . Also  $L_{\phi} \gg \xi$   $L_{\phi} = \xi \exp \left[ \pi \sqrt{\frac{\hbar n_1}{2m a \omega_{\perp}}} \right]$

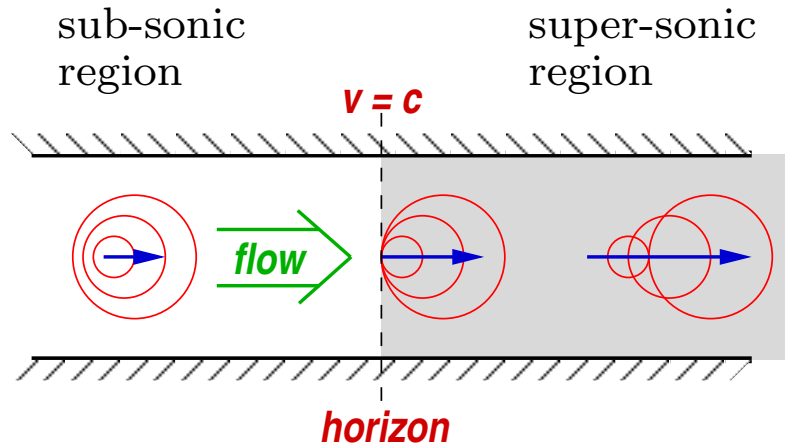
• the second inequality allows to avoid the 3D-like **transverse Thomas-Fermi regime** and implies that transverse motion is frozen



C. Tozzo & F. Dalfovo, Phys. Rev. A (2002)

←  $\eta = \mu/\hbar\omega_{\perp}$   
only  
axi-symmetric ex-  
citations  
included ( $m = 0$ )

# Sonic black holes : “dumb holes”



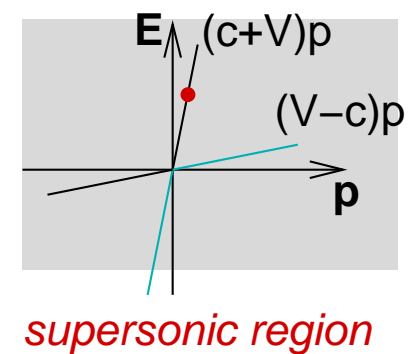
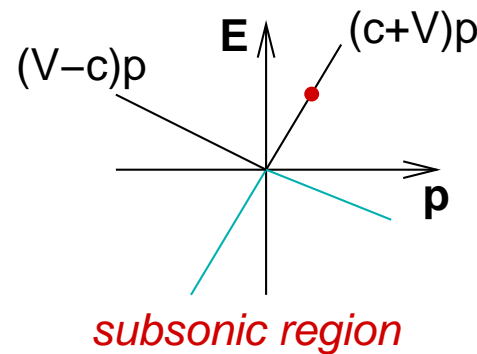
W. G. Unruh, Phys. Rev. Lett. (1981)

even without a source,  
vacuum fluctuations  $\rightsquigarrow$   
Hawking radiation

in the laboratory :

$$E(p) = c|p| + Vp$$

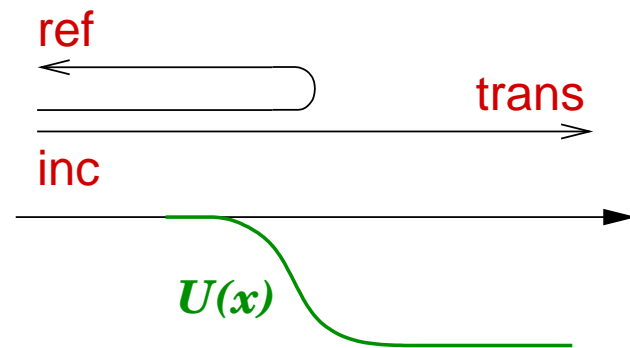
$\swarrow$                        $\searrow$   
 comoving                  Doppler



Analogous to tunnel effect : (quantum reflexion)

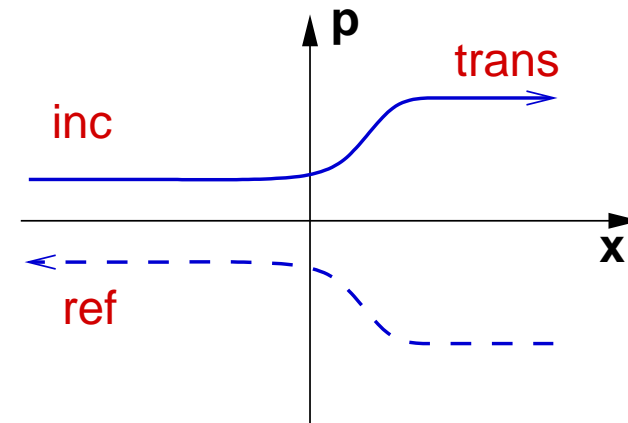
real space

particle incoming from the left with  $E > U_{\max}$

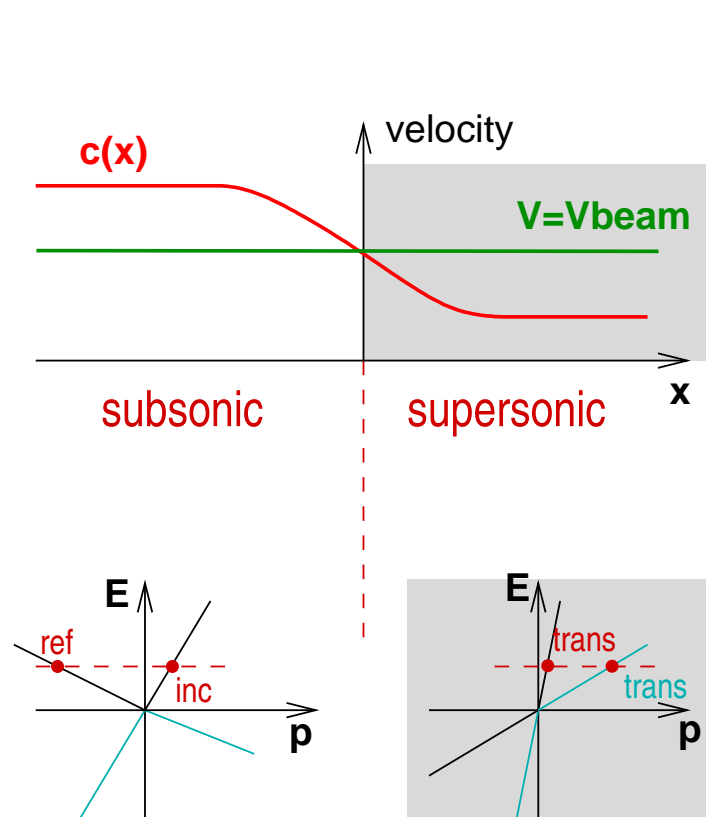


phase space trajectory

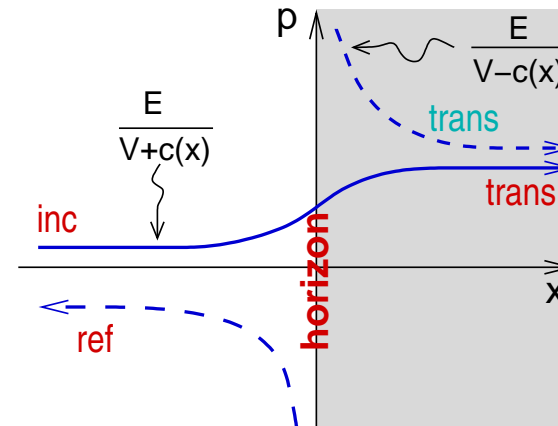
$$E = p^2/2m + U(x)$$



A model configuration :



$$H(x, p) = c(x) |p| + V p$$

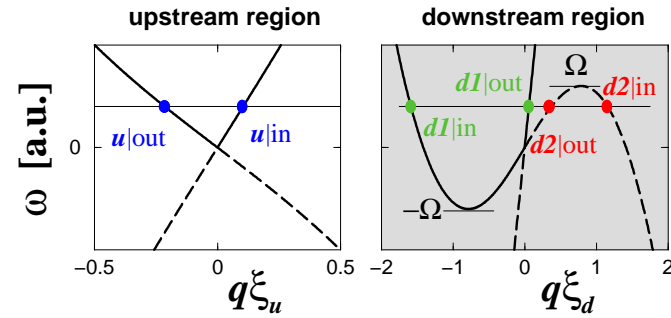
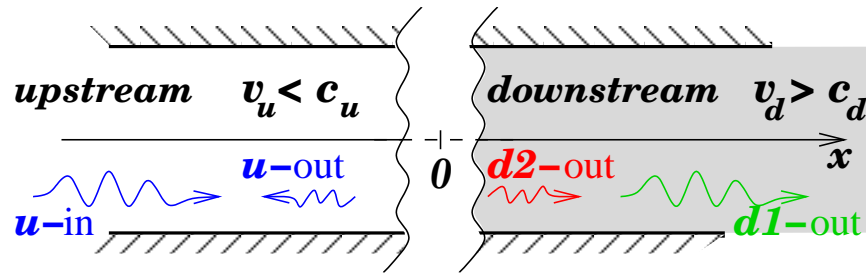


$$\text{tunnel proba } R \propto \exp \left\{ -\frac{2S}{\hbar} \right\}$$

$$S = \left| \text{Im} \int p(x) dx \right| \simeq \frac{\pi E}{c'(0)}$$

of the form  $R \propto \exp \{ -E/k_B T_H \}$   
with  $T_H \sim 10$  nK very weak ...

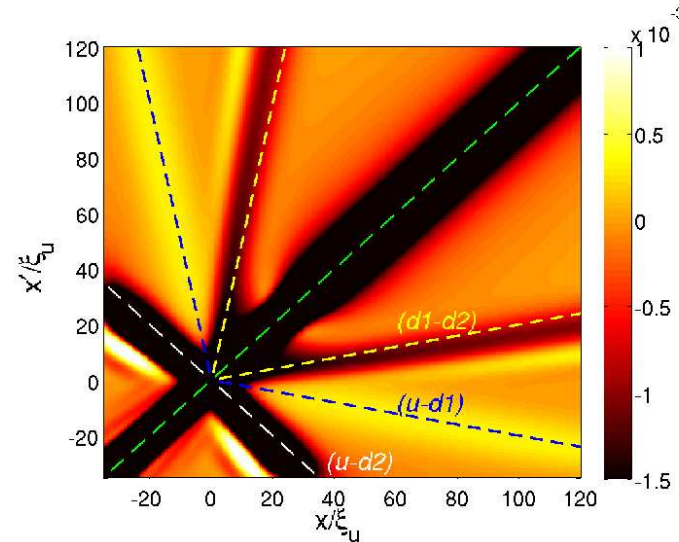




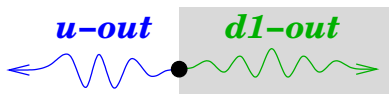
new theoretical and experimental interest :  
 study of density correlation on each side  
 of the horizon

$$G^{(2)}(x, x') = \frac{\langle : n(x)n(x') : \rangle}{\langle n(x') \rangle \langle n(x) \rangle} - 1$$

Balbinot, Carusotto, Fabbri, Fagnocchi, Recati  
 Phys. Rev. A (2008) & New J. Phys. (2008)



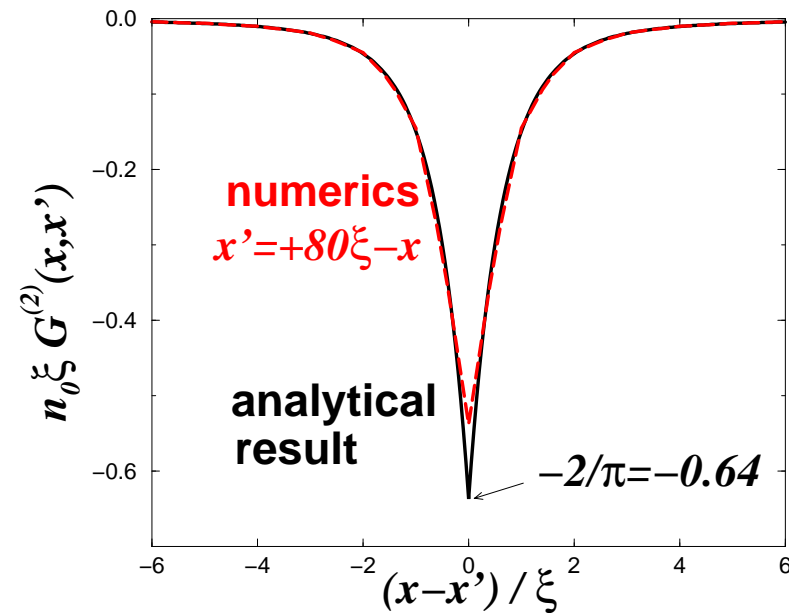
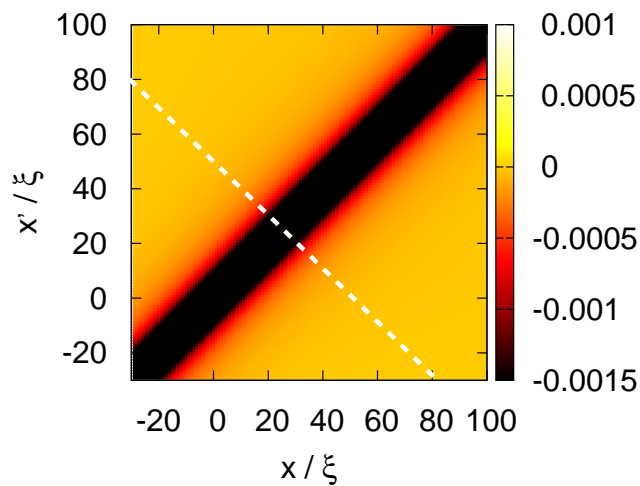
example :



$$x = (v_d + c_d)t \quad \text{correlates with} \quad x' = (v_u - c_u)t$$

## Uniform 1D Condensate (no black hole)

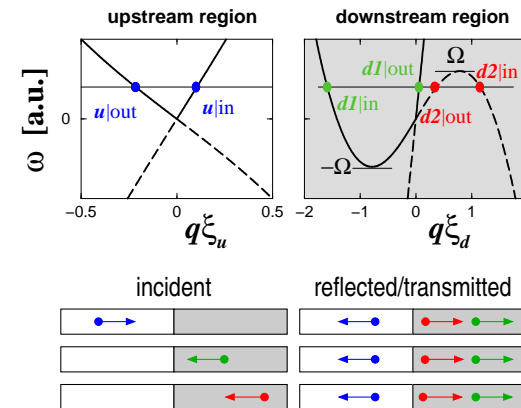
$$G^{(2)}(x, x') = \frac{-2}{\pi n_0 \xi} F\left(\frac{|x - x'|}{\xi}\right) \quad \text{where} \quad F(X) = \frac{1}{2X} \int_0^\infty dt \frac{\sin(2tX)}{(1+t^2)^{3/2}}.$$



In practice :

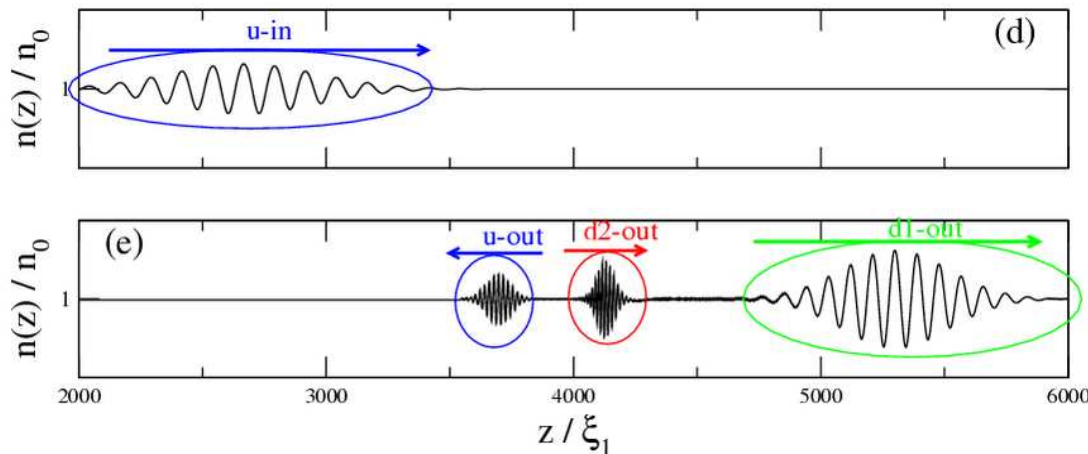
$$(\omega - k V_{\text{beam}})^2 = \omega_B^2(k)$$

$$\hat{\psi} = \psi_0 + \delta\hat{\psi} \text{ with } \delta\hat{\psi}(x) = \sum(3 \text{ modes})$$



Model configuration :  $U(x)$  and  $g(x)$  step like with

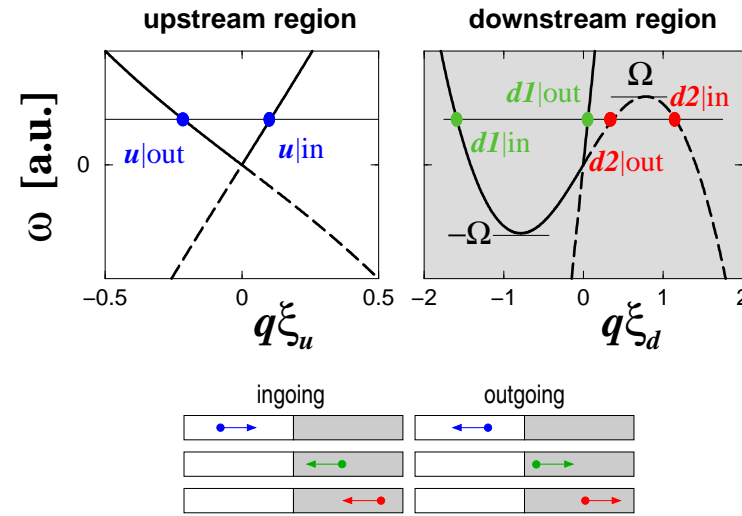
$$U(x) + g(x)n_0 = C^{\text{st}} \text{ such that } \psi_0(x) = \sqrt{n_0} \exp\{ik_0x\}, \forall x.$$



## One-body Hawking signal

linear relation connecting the operators of the out-going modes  $\hat{b}_{u,d1,d2}$  to the in-going  $\hat{a}_{u,d1,d2}$  ones

$$\begin{pmatrix} \hat{b}_u(\omega) \\ \hat{b}_{d1}(\omega) \\ \hat{b}_{d2}^\dagger(\omega) \end{pmatrix} = \mathbf{S}(\omega) \begin{pmatrix} \hat{a}_u(\omega) \\ \hat{a}_{d1}(\omega) \\ \hat{a}_{d2}^\dagger(\omega) \end{pmatrix}.$$



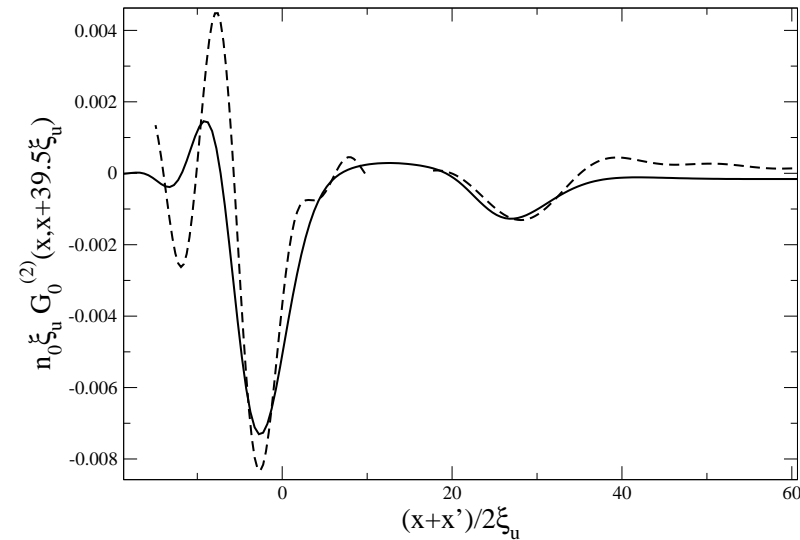
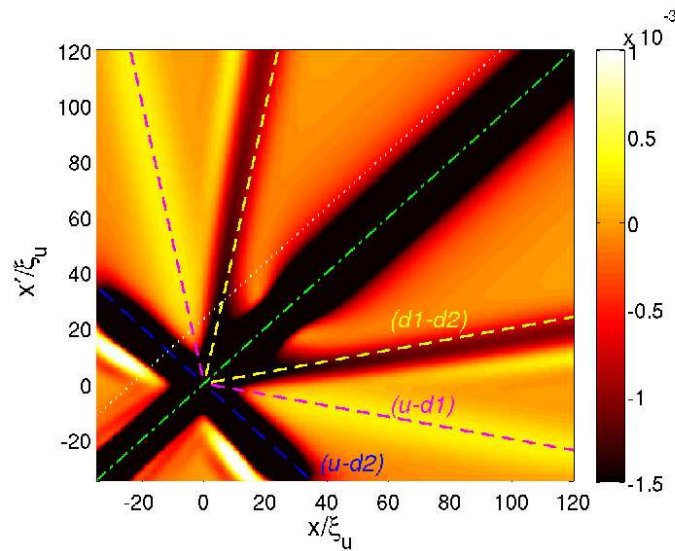
Radiation in the subsonic region occurs in the  $u$ -outgoing mode with

$$\frac{dI_u^{\text{out}}}{dt d\omega} = \langle \hat{b}_u^\dagger(\omega) \hat{b}_u(\omega) \rangle = |\mathbf{S}_{uu}|^2 I_u^{\text{in}} + |\mathbf{S}_{ud1}|^2 I_{d1}^{\text{in}} + |\mathbf{S}_{ud2}|^2 (I_{d2}^{\text{in}} + 1).$$

at  $T = 0$  :  $\frac{dI_u^{\text{out}}}{dt d\omega} = |\mathbf{S}_{ud2}|^2$  needs  $\left\{ \begin{array}{l} u \rightleftharpoons d2 \text{ mode conversion} \\ d2\text{-ingoing mode !} \end{array} \right.$

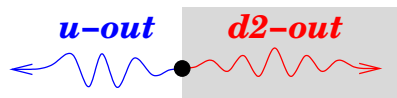
## Two-body Hawking signal

Comparison of numerical and analytic results (stationary phase neglecting interferences between the correlation signals) :



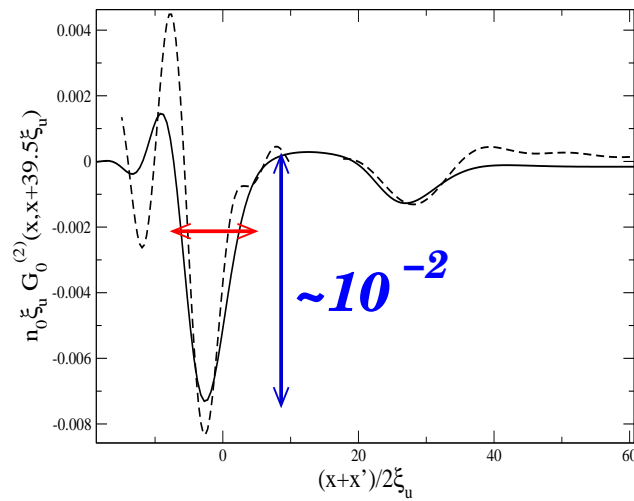
A. Recati, N. Pavloff & I. Carusotto, Phys. Rev. A **80**, 043603 (2009)

main correlation signal :



$$x = V_{d2-out} t \quad \text{correlates with} \quad x' = V_{u-out} t$$

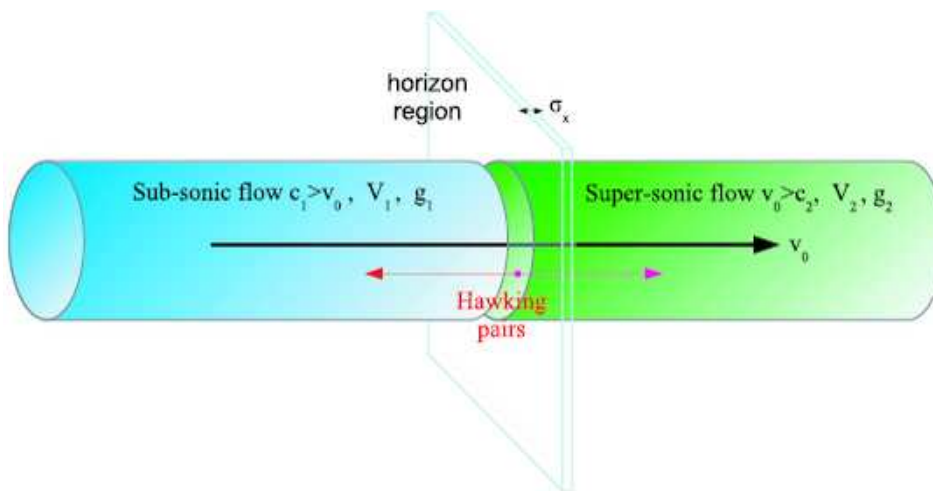
orders of magnitude :



<sup>87</sup>Rb

	laser	cigar
$n_0 \sim$	$50 \mu\text{m}^{-1}$	$100 \mu\text{m}^{-1}$
$c \sim$	$1 \text{ mm/s}$	$3 \text{ mm/s}$
$\xi \sim$	$1 \mu\text{m}$	$0.1 \mu\text{m}$

$$\left| G^{(2)} \right|_{\text{max}} \sim 2 \times 10^{-4} \leftrightarrow 10^{-3}$$



need for new dumb-hole configurations !

## Conclusion

Density correlations appear as promising tools for identifying Hawking radiation ...with some **un**essential limitations.

→ **Clear signal** , well understood. One knows where to look, and at which quantity.

→ **Poorly affected by noise and finite  $T$**  .

→ **Drawbacks** :  
| weak signal intensity.  
| awkward configuration.  
| what about transverse degrees of freedom ?

→ **What comes next ?** | more realistic dumb hole configurations,  
| white hole stability ...

## Conclusion (bis)

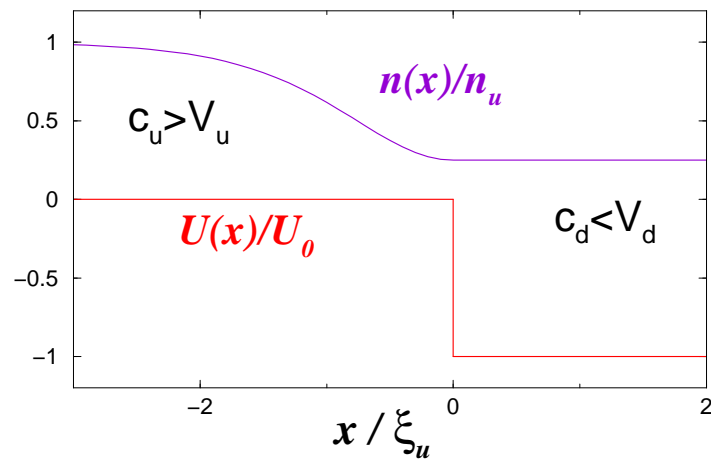
General comments concerning some of the physical problems that can be investigated by means of atom lasers.

- **transport in presence of disorder** . Effects of **interaction** lead to qualitatively different phenomena. What is the transmission in the time dependent regime ?
- **Hawking radiation** , BECs seem to offer the most promising prospect to observe a **fully quantum** Hawking radiation.
- **Dispersive shocks** , BEC appears to be a versatile tool for studying **frictionless** nonlinear mechanisms of dissipation.

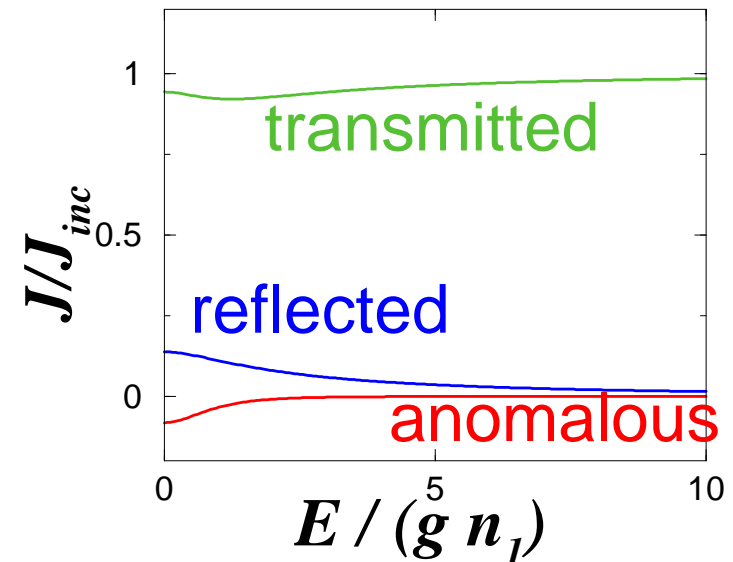
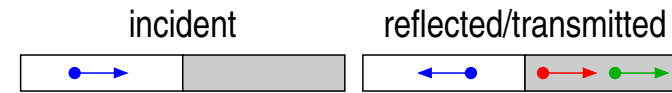


## Waterfall configuration

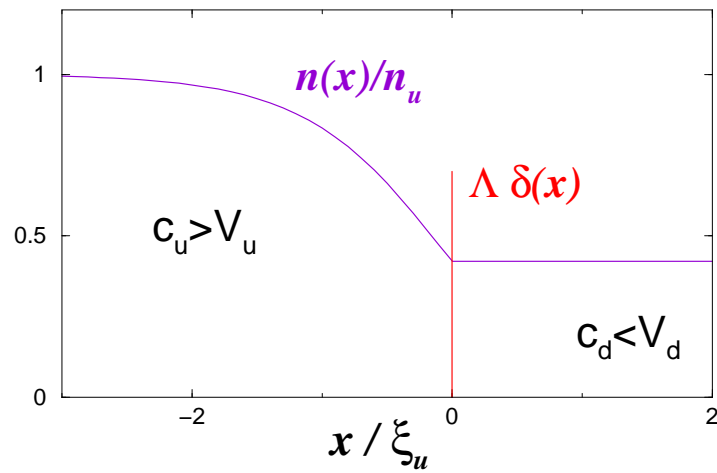
for instance :



$$V_d/c_d = 0.25 \quad V_d/c_d = 16$$

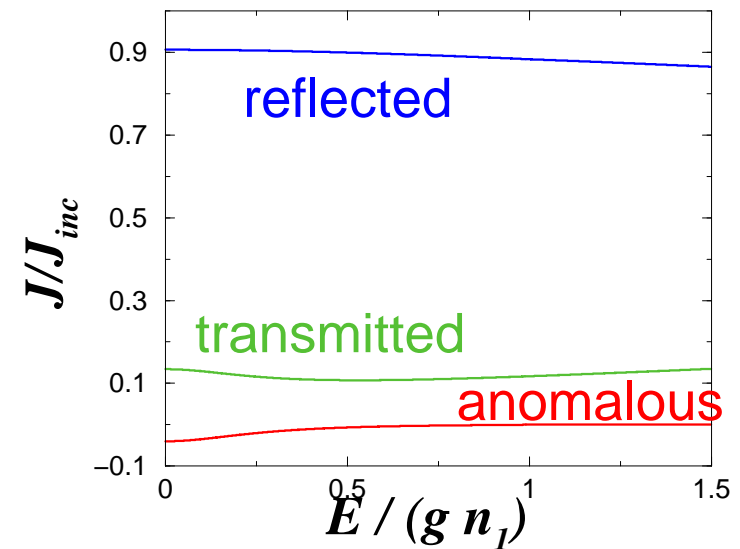


## Localized obstacle



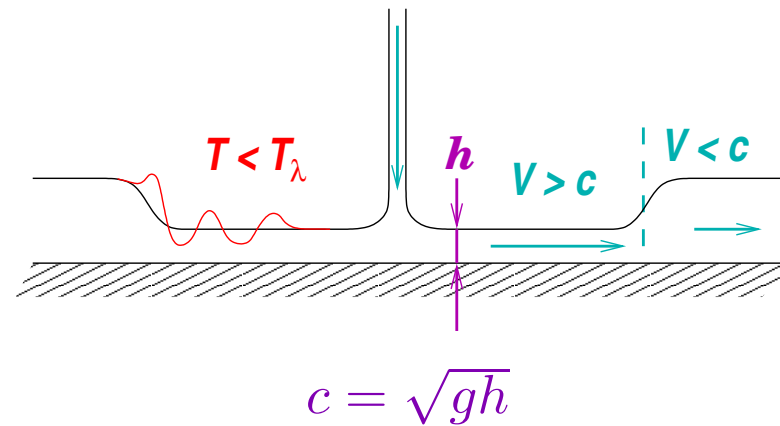
$$V_u/c_u = 0.1 \quad V_d/c_d = 5.0$$

for instance :



## At longer term ...

The hydraulic jump is a stable white hole (Volovik JETP 2005)



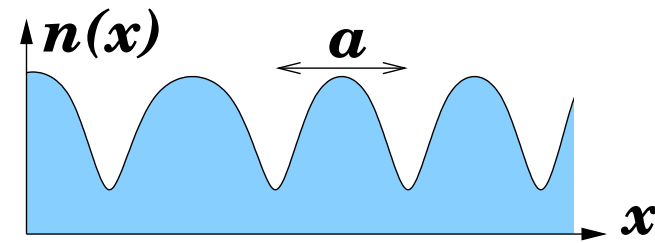
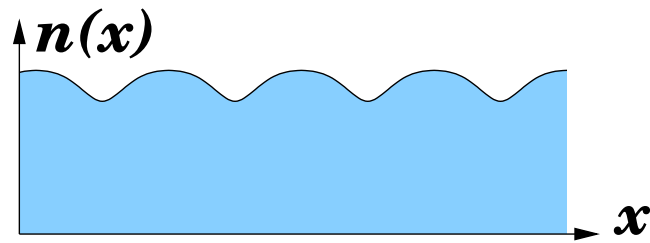
$$(\omega - \vec{k} \cdot \vec{v})^2 = c^2 k^2 \left[ 1 + h^2 k^2 \left( -\frac{1}{3} + \frac{\sigma}{\rho g h^2} \right) + \dots \right]$$

appearance of oscillations in the superfluid phase ?

cf, Pitaevskii striped phase ?

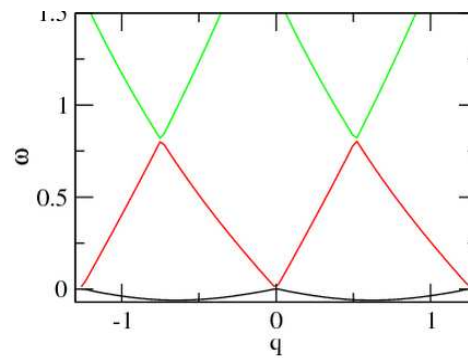
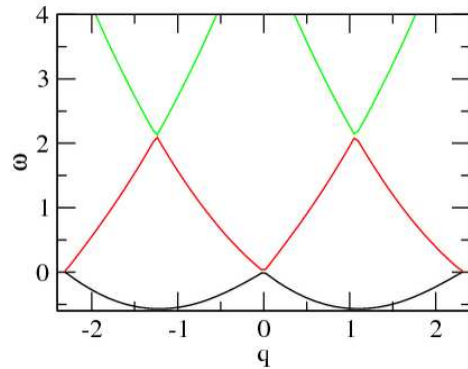
## 1D Super-solid

L. Pitaevskii (JETP 84): above the Landau critical velocity, a super-sonic superfluid forms a “striped phase”



**Question:** is this “supersolid” phase superfluid ?

One has to study the excitation spectrum :



$$\omega \simeq -\sqrt{\frac{k}{m}} \left| \sin\left(\frac{qa}{2}\right) \right|$$

